# Preparations for Measuring the Gamma Ray Strength Function of <sup>60</sup>Fe using <sup>59</sup>Fe(d,p)<sup>60</sup>Fe

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### Introduction

- <sup>60</sup>Fe has been observed nearby in the galaxy <sup>[1]</sup>, in lunar soil samples <sup>[2]</sup>, in oceans worldwide <sup>[3]</sup>, and in Antarctic snow<sup>[4]</sup>
- Important isotope in nucleosynthesis; branch point in the s-process
- We will measure the gamma ray strength function in order to obtain a constraint for astrophysical models
- <sup>59</sup>Fe beam, produced using the TAMU MARS line at the Cyclotron Institute, impinging on a CD<sub>2</sub> target
- Proton emitted at backward angles from <sup>59</sup>Fe(d,p)<sup>60</sup>Fe reaction will be detected in a silicon detector
- Gamma rays will be detected in packs of BaF<sub>2</sub> crystals coupled to photomultiplier tubes.

#### **Experimental Setup**





CAD drawing of anticipated final experimental setup

- Coincidence of MCP and PPAC give TOF of residue
- Silicon rings give angular resolution for proton
- Purity detector for beam impurities event by event • ~10<sup>5</sup> pps



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## **BaF**<sub>2</sub> Resolution Improvement

- Decoupled to try RTV
- Initially successful but found degradation over time and uniformity issues
- Recover fast component of signal when recoupled with
- Resolution still between 15-25%
- New PMTs increase resolution of one detector from 17.7% to 11.6%







### **Coming This Fall**

- Measure <sup>57</sup>Fe(d,p)<sup>58</sup>Fe reaction in order to characterize detector response
- Can check with known states of <sup>58</sup>Fe

#### References

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